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LETTER TO THE EDITOR

Integrable systems admitting topological solitons

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Abstract. We discuss $(1 + 1)$ -dimensional sigma models and Heisenberg models in which the target space has the topology of a cylinder. In these integrable systems, the solutions are classified by a winding number.

1. Introduction

This letter deals with integrable $(1 + 1)$ -dimensional systems, in which the configuration space or phase space has disconnected sectors classified topologically by an integer (a winding number). The general set-up involves a field $\Phi(x, t)$ taking values in some manifold M . Here $t \in \mathbb{R}$ denotes time, and $x \in X$ is the space variable; either $X = \mathbb{R}$ and we impose a boundary condition $\Phi(-\infty, t) = \Phi(\infty, t)$, or $X = S^1$ (i.e. Φ is periodic in x). Furthermore, M is chosen so that its fundamental group equals the group of integers \mathbb{Z} . So for each fixed t , $\Phi(\cdot, t)$ is in effect a continuous mapping from a circle into M , and it therefore has a winding number $N \in \mathbb{Z}$. As t changes, this integer remains constant. One could also view a solution Φ as wrapping the spacetime $X \times \mathbb{R}$ around the image manifold M , with this mapping having winding number N (in an appropriate sense).

The prototype in the $X = \mathbb{R}$ case is the sine-Gordon equation, where $M = S^1$. If we think of Φ as a two-dimensional unit vector ($\Phi \cdot \Phi = 1$), then the sine-Gordon system corresponds to the Lagrangian density

$$\mathcal{L} = \frac{1}{2} \eta^{\mu\nu} \Phi_\mu \cdot \Phi_\nu - (1 - K \cdot \Phi) \quad (1)$$

where K is a constant unit vector. Hereafter $x^\mu = (x^0, x^1) = (t, x)$ are the spacetime coordinates; a subscript denotes partial differentiation; and $\eta^{\mu\nu} = \text{diag}(1, -1)$ is the (inverse) spacetime metric. An N -kink solution of the sine-Gordon equation has winding number N .

In our examples M will be the cylinder $S^1 \times \mathbb{R}$, which (at least to begin with) we think of as the hyperboloid of one sheet in \mathbb{R}^3 . One can visualize the field as a closed string which is wound around this cylinder, and evolves in time. Instead of Φ , let us use the symbol ψ^a , denoting a three-dimensional vector satisfying

$$\eta_{ab} \psi^a \psi^b = 1 \quad (2)$$

where $\eta_{ab} = \text{diag}(1, 1, -1)$. The metric on the hyperboloid (2) is taken to be the one induced by the metric η_{ab} ; M is then a symmetric space $SO(2, 1)/SO(1, 1)$. Two naturally defined systems taking values on M are the nonlinear sigma model, and the Heisenberg model (Landau–Lifshitz equation). The σ -model is defined by the Lagrangian density

$$\mathcal{L} = \frac{1}{2} \eta^{\mu\nu} \psi_\mu^a \psi_\nu^b \eta_{ab} \quad (3)$$

while the Landau–Lifshitz equations arise from the Hamiltonian density

$$\mathcal{H} = \frac{1}{2} \psi_x^a \psi_x^b \eta_{ab} \tag{4}$$

with Poisson brackets

$$\{\psi^a(x), \psi^b(y)\} = -\delta(x - y) \varepsilon^{abc} \psi^d(x) \eta_{cd}. \tag{5}$$

The analogous S^2 systems, in which η_{ab} is replaced by the Euclidean metric δ_{ab} , are of course well known integrable systems; and their hyperbolic versions are therefore integrable too, simply by analytic continuation. (By integrability we mean the existence of a suitable Lax pair.) In fact, these hyperbolic versions have been widely studied in their own right: cf [1–6]. The aim in what follows is to examine the simplest ‘winding’ solutions of the systems (3) and (4), (5).

2. The hyperbolic Heisenberg model

The Landau–Lifshitz equation obtained from (4) and (5) is

$$\psi_t^a = \eta^{ab} \varepsilon_{bcd} \psi^c \psi_{xx}^d. \tag{6}$$

If we parametrize the hyperboloid in terms of ‘polar angles’ as

$$\psi^a = (\cosh \theta \cos \phi, \cosh \theta \sin \phi, \sinh \theta)$$

then (6) is equivalent to

$$\theta_t = 2(\sinh \theta) \theta_x \phi_x + (\cosh \theta) \phi_{xx} \tag{7}$$

$$\phi_t = (\operatorname{sech} \theta) \theta_{xx} + (\sinh \theta) (\phi_x)^2 \tag{8}$$

while if we parametrize in terms of a stereographic projection as

$$\psi^a = \frac{1}{1 + u^2 - v^2} (1 - u^2 + v^2, 2u, 2v)$$

then (6) becomes

$$u_t = -v_{xx} - 2\alpha v(u_x^2 + v_x^2) + 4\alpha u u_x v_x \tag{9}$$

$$v_t = -u_{xx} + 2\alpha u(u_x^2 + v_x^2) - 4\alpha v u_x v_x \tag{10}$$

where $\alpha = (1 + u^2 - v^2)^{-1}$.

Let us first look for static winding solutions, by solving (7), (8) with θ and ϕ being functions of x only. From (7) we obtain

$$\phi_x = B \operatorname{sech}^2 \theta \tag{11}$$

where B is a constant; and then (8) integrates to

$$\theta_x = \sqrt{A + B^2 \operatorname{sech}^2 \theta} \tag{12}$$

with A constant. For a winding solution, we want $\theta(x)$ to be non-monotonic, which requires $-B^2 \leq A < 0$. Write $A = -N^2$, where $0 < N \leq B$. Then the solution of (12) is

$$\theta(x) = \sinh^{-1} \left(\sqrt{B^2/N^2 - 1} \sin(Nx) \right) \tag{13}$$

and this is periodic, with period 2π , provided N is an integer. Finally, $\phi(x)$ is obtained by integrating the smooth function (11). To check that we have a winding solution, it is sufficient to compute

$$\Delta\phi = \int_0^{2\pi} B \operatorname{sech}^2 \theta \, dx \tag{14}$$

we find that $\Delta\phi = 2\pi N$, and so we have a space-periodic static solution, with winding number N .

Note that if $B = N$ in (13), then $\theta = 0$ and $\phi = Nx$: a solution which winds N times around the ‘waist’ of the hyperboloid. The simplest time-dependent solution is a generalization of this, namely $\theta \equiv \text{constant}$, $\phi = Nx + N^2(\sinh \theta)t$.

For these solutions, the space X is the circle; in fact, there are no travelling-wave winding solutions for $X = \mathbb{R}$. One can find a simple time-dependent solution on \mathbb{R} by using the stereographic form (9), (10) of the equations. Let us look for solutions in which α is a function of x only, i.e. take $u^2 - v^2 = f(x)^2$. It follows from (9) and (10) that the function $g = vu_x - uv_x$ satisfies

$$\frac{\partial g}{\partial x} = 2g \frac{\partial}{\partial x} \log(1 + f^2).$$

If we take the simplest solution of this, namely $g \equiv 0$, then u and v must have the form

$$\begin{aligned} u(x, t) &= f(x) \cosh h(t) \\ v(x, t) &= f(x) \sinh h(t). \end{aligned}$$

Substituting these into (9), (10) gives $dh/dt = -m$ constant, and

$$\frac{d^2 f}{dx^2} = \frac{2f}{(1 + f^2)} \left(\frac{df}{dx} \right)^2 + mf. \quad (15)$$

This has the first integral

$$\left(\frac{df}{dx} \right)^2 = c(1 + f^2)^2 - m(1 + f^2)$$

where c is an arbitrary constant. For winding solutions one needs $c \geq 0$. If $c = 0$, one obtains a solution which is equivalent to (18) below; so take $c > 0$, and scale x so that $c = 1$. Now the requirement of winding imposes a restriction on m , namely $m \leq 1$. For $m < 0$ one obtains solutions which are similar to those for $m \geq 0$, so let us take $0 \leq m \leq 1$. Then the solution of (15) is

$$f(x) = \sqrt{1 - m} \operatorname{sc}(x|m) \quad (16)$$

in the elliptic-function notation of [7]. The two limits of (16) are

$$f(x) = \tan x \quad \text{for } m = 0 \quad (17)$$

and (after a shift in x)

$$f(x) = \operatorname{cosech} x \quad \text{for } m = 1. \quad (18)$$

Therefore, we have a family of winding solutions parametrized by $m \in [0, 1]$, with $h(t) = -mt$ and where $f(x)$ is specified in (16), (17) or (18). For (16) and (17), the solution is periodic (i.e. $X = S^1$); while for (18) it lives on $X = \mathbb{R}$.

A Lax pair corresponding to the equation (6) is as follows. Define a 2×2 matrix $S \in SL(2, \mathbb{R})$ by

$$S = \begin{bmatrix} \psi^1 & \psi^2 + \psi^3 \\ \psi^2 - \psi^3 & -\psi^1 \end{bmatrix}.$$

Then the consistency condition for the linear system

$$\begin{aligned} \Psi_x &= \lambda S \Psi \\ \Psi_t &= -\lambda(2\lambda S + S_x S) \Psi \end{aligned}$$

is exactly (6). Implementation of the inverse scattering transform (on $X = \mathbb{R}$) produces N -soliton solutions, and these turn out to have winding number N . The solution corresponding to (18), which in terms of ψ^a is

$$\psi^a = (1 - 2 \operatorname{sech}^2 x, 2 \operatorname{sech}^2 x \sinh x \cosh t, -2 \operatorname{sech}^2 x \sinh x \sinh t)$$

is the simplest example: a stationary 1-soliton with winding number $N = 1$. These solitons are closely related to those of the ‘standard’ Heisenberg model [8].

3. The hyperbolic sigma model

In terms of the (θ, ϕ) parametrization, the σ -model equations are

$$\theta_{tt} - \theta_{xx} = -\cosh \theta \sinh \theta (\phi_t^2 - \phi_x^2) \quad (19)$$

$$(\phi_t \cosh^2 \theta)_t = (\phi_x \cosh^2 \theta)_x. \quad (20)$$

For the static problem, the equations (and hence their solutions) are the same as for the static Heisenberg case of the previous section. What follows are some examples of time-dependent winding solutions.

First, let us look for solutions on $X = S^1$ for which $\phi = x$. (To go to winding number $N > 1$ is an easy generalization.) It follows from (19), (20) that θ is a function of t only, satisfying

$$\theta_t^2 = c + \sinh^2 \theta$$

where c is an arbitrary constant. This is easily integrated in terms of elliptic functions. Apart from the case $c = 0$, the solutions have the property that $\theta(t)$ reaches infinity in finite time. An example is $c = 1$, where the solution

$$\theta(t) = \sinh^{-1} \tan t$$

goes from $\theta = -\infty$ to $+\infty$ as t goes from $-\pi/2$ to $+\pi/2$. For $c > 0$ the solutions all have this behaviour, whereas for $c < 0$, θ comes in from infinity, turns round, and goes out again. In the limiting case $c = 0$, the solution

$$\theta(t) = 2 \tanh^{-1}(ke^t)$$

includes the static case $\theta \equiv 0$, and tends asymptotically to $\theta = 0$ as $t \rightarrow -\infty$.

Our second family of solutions arises from the ‘self-duality’ equations

$$\phi_x = (\operatorname{sech} \theta)\theta_t$$

$$\phi_t = (\operatorname{sech} \theta)\theta_x$$

which imply (19), (20). So ϕ and $\mu = 2 \tan^{-1} \exp \theta$ satisfy $\phi_x = \mu_t$, $\phi_t = \mu_x$, and are therefore ‘conjugate’ solutions of the (1 + 1)-dimensional wave equation. The general solution is

$$\phi = f(x + t) + g(x - t)$$

$$\mu = f(x + t) - g(x - t)$$

where f and g are arbitrary functions. All winding solutions are defined only on a finite time interval. For example, the choice

$$f(\xi) - \frac{\pi}{4} = g(\xi) + \frac{\pi}{4} = \frac{1}{2}\pi \tanh \xi$$

leads to

$$\begin{aligned}\phi &= \frac{\pi \sinh 2x}{\cosh 2x + \cosh 2t} \\ \mu &= \frac{\pi}{2} + \frac{\pi \sinh 2t}{\cosh 2x + \cosh 2t}\end{aligned}$$

which is a winding solution on \mathbb{R} (with unit winding number). However, bearing in mind that we need $0 < \mu < \pi$, we see that it represents a smooth solution only for $|t| < \frac{1}{2} \log 3$.

4. Positive-definite versions

In the systems discussed above, the metric on the target space M is indefinite:

$$\begin{aligned}ds^2 &= \eta_{ab} d\psi^a d\psi^b \\ &= -d\theta^2 + \cosh^2 \theta d\phi^2.\end{aligned}$$

One can replace this by an analogous positive-definite metric on the cylinder, namely

$$ds^2 = d\theta^2 + \cosh^2 \theta d\phi^2. \quad (21)$$

The corresponding equations remain integrable, since they are obtained by simply making the replacement $\phi \mapsto i\phi$ (and, for the Landau–Lifshitz case, also $t \mapsto -it$). We are still thinking of ϕ as being a periodic coordinate (and looking for solutions which wind in ϕ); as a consequence of this, M is no longer a symmetric space.

Let us briefly look at the corresponding ‘sigma model’. From the Lagrangian density

$$\mathcal{L} = \frac{1}{2} \eta^{\mu\nu} [(\cosh^2 \theta) \phi_{,\mu} \phi_{,\nu} + \theta_{,\mu} \theta_{,\nu}]$$

one obtains the equations of motion

$$\begin{aligned}\theta_{tt} - \theta_{xx} &= \cosh \theta \sinh \theta (\phi_t^2 - \phi_x^2) \\ (\phi_t \cosh^2 \theta)_t &= (\phi_x \cosh^2 \theta)_x.\end{aligned}$$

The most general static winding solution is now $\theta \equiv 0$, $\phi = Nx$. This is what one would expect: in the positive-definite case, the string will try to minimize its length, and $\theta = 0$ is where the cylinder is narrowest (with respect to the metric (21)). If we look for more general solutions having $\phi = Nx$, then θ has to be a function of t only, with

$$\tanh \theta = \sqrt{m} \operatorname{sn}(\rho t | m)$$

where ρ and m are constants with $\rho > |N|$ and $m = 1 - N^2/\rho^2$. In other words, the string oscillates between the values $\theta_{\pm} = \pm \tanh^{-1} \sqrt{m}$.

5. Concluding remarks

There are several examples of integrable elliptic systems of partial differential equations admitting topological soliton solutions: for example, instantons in sigma models on \mathbb{R}^2 and gauge theory on \mathbb{R}^4 , and BPS monopoles on \mathbb{R}^3 . In these cases, there is no time dependence.

Analogous time-dependent examples (in other words, hyperbolic or parabolic systems, rather than elliptic) are not as prevalent: in fact, sine-Gordon is the only well known example. But many such integrable systems exist, and in this note we have briefly examined a few of them. It might be of interest to study further the inverse scattering transforms for these cases, and to try to attempt a general classification of systems of this type.

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